

# Can we use current/past instruments to get us part way there?

Instrument	IRIS	IMG	AIRS	IASI
Satellite	Nimbus 4	ADEOS	AQUA	METOP-A
Spectro- meter type	FTS	FTS	grating spectrometer	FTS
Data available	Apr 1970 – Jan 1971	Oct 1996 – Jun 1997	2002 - present	2007 - present
Spectral coverage (cm <sup>-1</sup> )	400 – 1600 cm <sup>-1</sup> continuous	715 – 3030 cm <sup>-1</sup> 3 bands	650 – 2700 cm <sup>-1</sup> 2378 bands	645 – 2760 cm <sup>-1</sup> 3 bands
Spectral resolution	2.8 cm <sup>-1</sup>	0.1 cm <sup>-1</sup>	0.4-1.0 cm <sup>-1</sup>	0.5 cm <sup>-1</sup>
Footprint (nadir)	95 km diameter	8km x 8km	13 km diameter	12 km diameter

Clear-sky only, no account of variability

Harries et al., 2001

→ Griggs and Harries, 2007 🗲

IASI and IRIS? IASI and IMG?

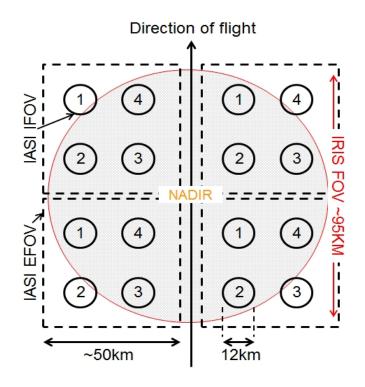
- What is the short-term variability seen in observed radiance spectra?
- How do these signals compare to those seen in model simulations and what can this tell us about the representation of the processes driving variability/change?
- Are observed long-term change signals robust?

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#### Ensure measurements are as consistent as possible

#### **Spatial consistency:**

average 16 IASI IFOV footprints



5 years of IASI L1c data: ~ 50 Tb ~ 160 million spectra (now archiving 2013-2014)

#### **Spectral consistency**

**IRIS** 

Pad each spectrum to 0-2000 cm<sup>-1</sup> at original sampling interval

FT padded spectrum

FT and output at 0.1 cm<sup>-1</sup> sampling interval (~ 2.8 cm<sup>-1</sup> resolution)

**IASI** 

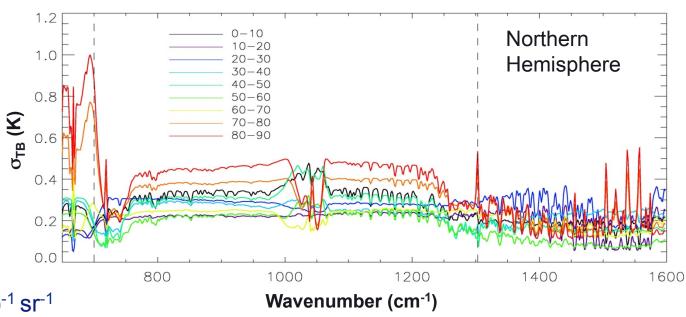
Pad and truncate average spectra to 0-2000 cm<sup>-1</sup> at original sampling interval

FT, remove IASI apodisation function & apply varying length Hamming window

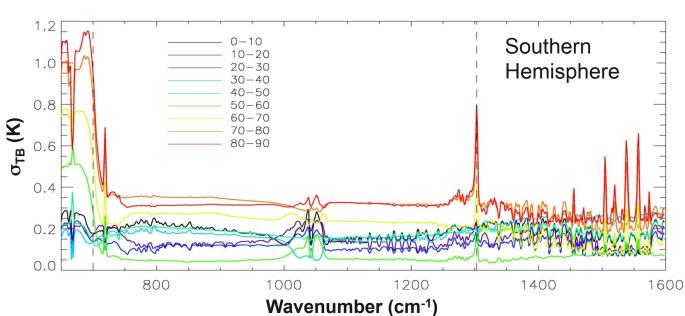
FT output at 0.1 cm<sup>-1</sup> sampling interval (~ 2.8 cm<sup>-1</sup> resolution)

Apply remaining FOV correction factor

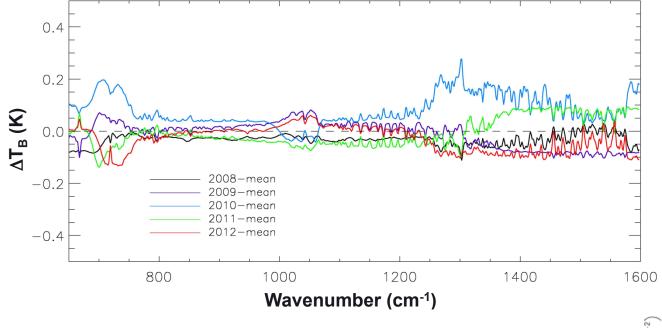
Standard deviation in 10° latitude band annual means



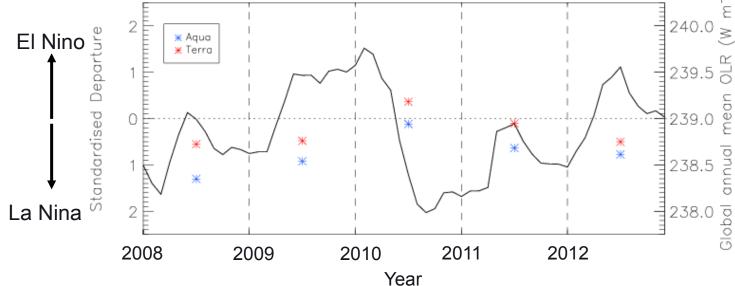
 $1 \text{ K} \sim 1 \text{ mW m}^{-2} (\text{cm}^{-1})^{-1} \text{sr}^{-1}$ 



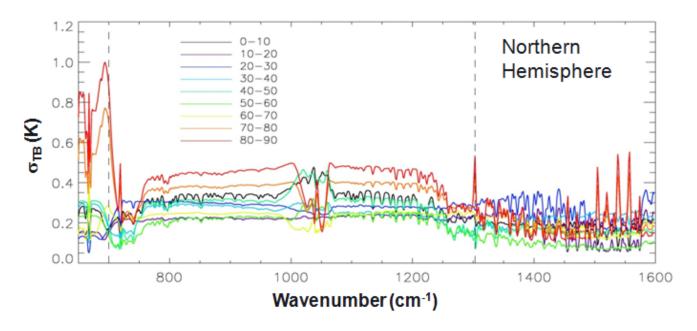
Deviation from overall global annual mean for each year



Multivariate
ENSO Index
(NOAA ESRL)
[black]
CERES OLR
[blue and red]

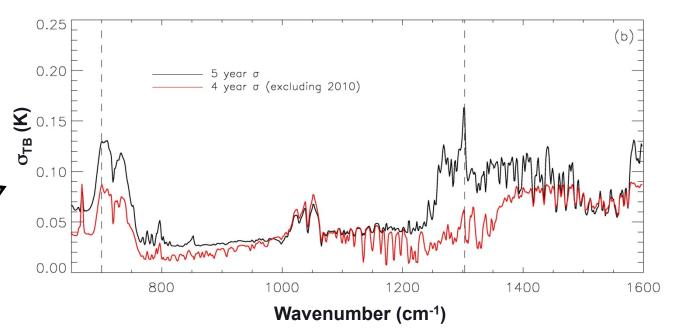


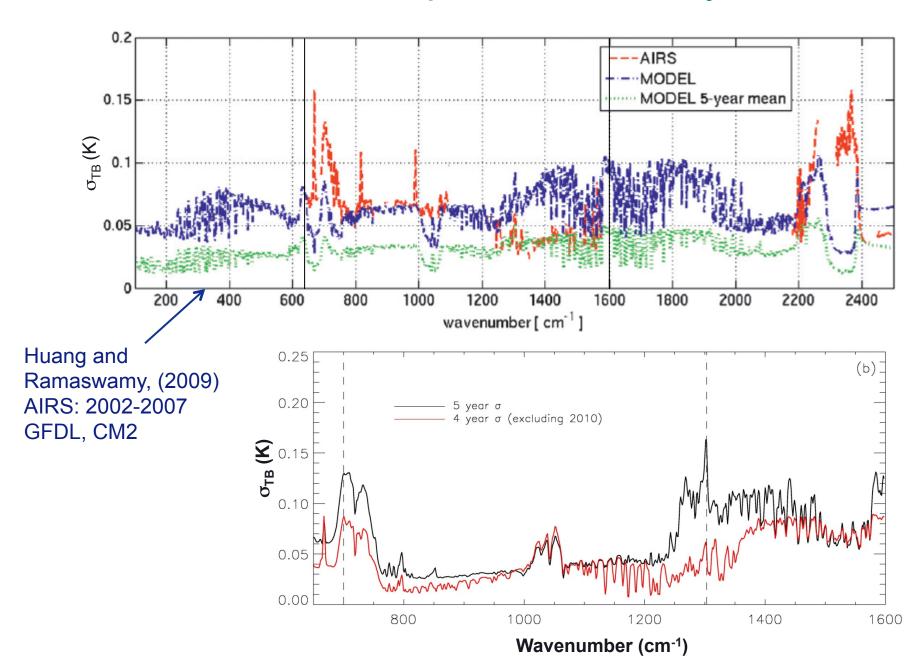
Standard deviation in 10° latitude band annual means





Note change in scale and change in shape

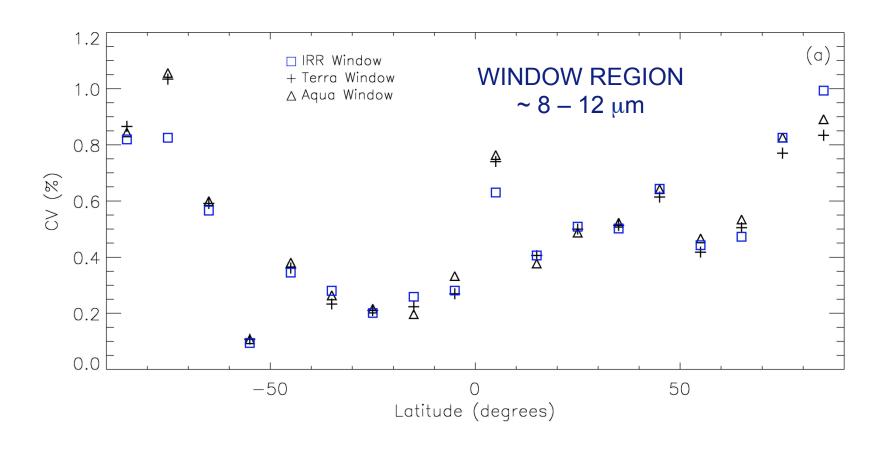




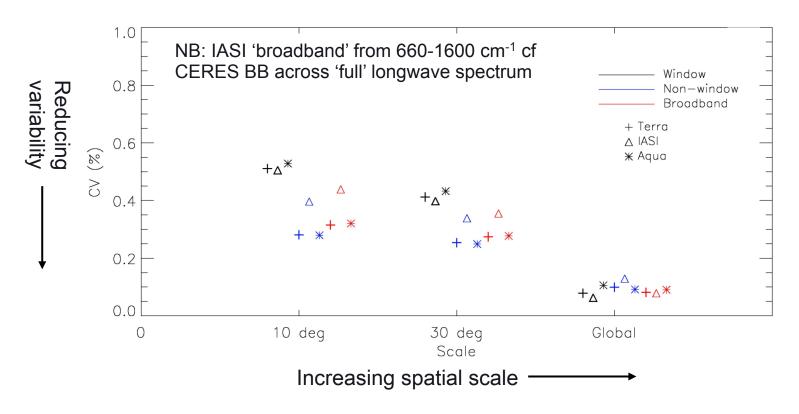
#### Consistency with broadband measurements?

Employ observations from CERES: broadband and window *fluxes*Different measurement scales so use *coefficient of variation, CV* 

CV = 
$$\sigma$$
 /  $\mu$  and note that  $\sigma_{BB} = [\sigma_{win}^2 + \sigma_{nonwin}^2 + 2cov_{win,nonwin}]^{1/2}$ 

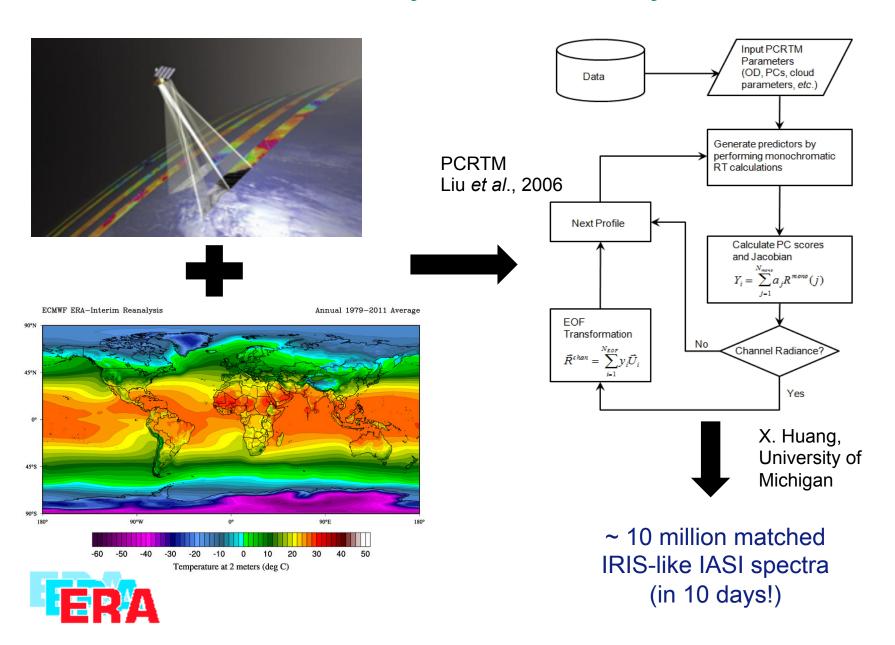


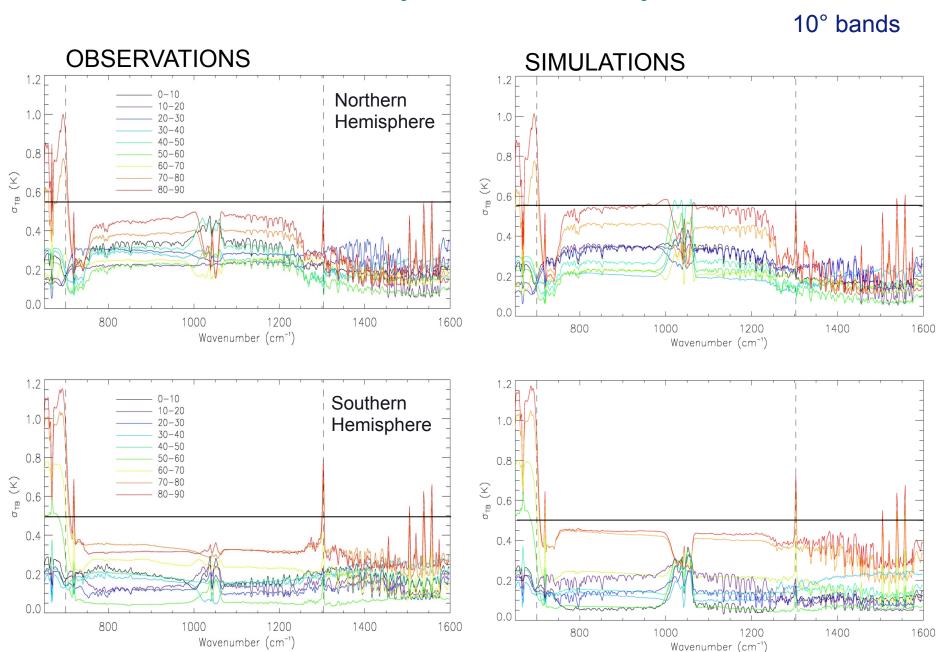
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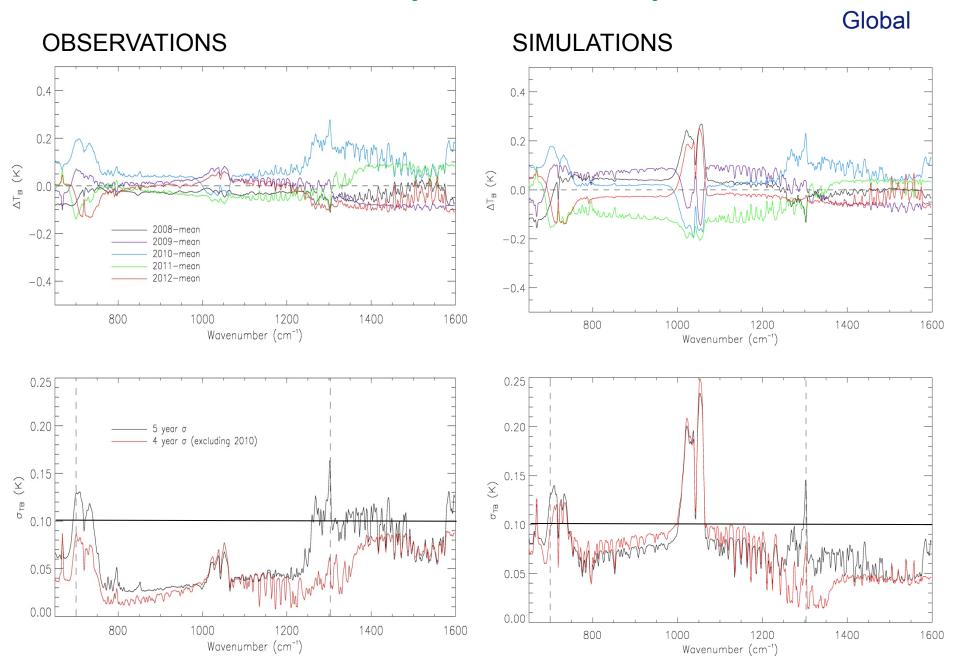


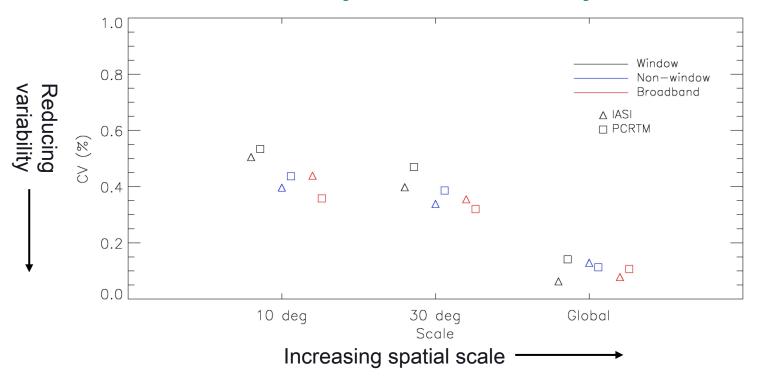
- Window inter-annual variability reduces most rapidly with increasing scale
- Results in non-window variability becoming dominant at global scale
- Difference between IASI BB and CERES BB behaviour suggests an important role for the far infra-red in determining all-sky inter-annual variability at the global scale
- Spectrally, global inter-annual variability < 0.17 K, < 0.05 K across window</li>

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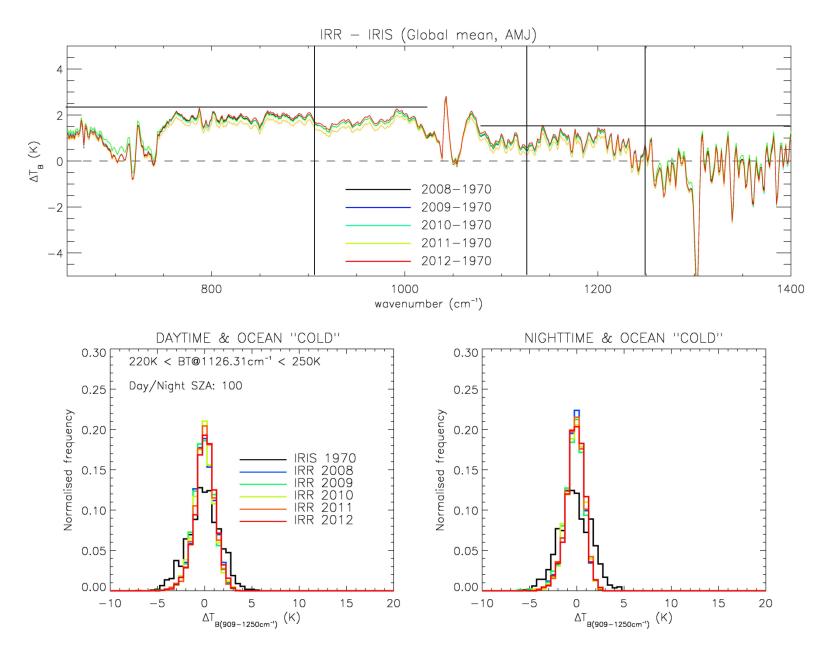




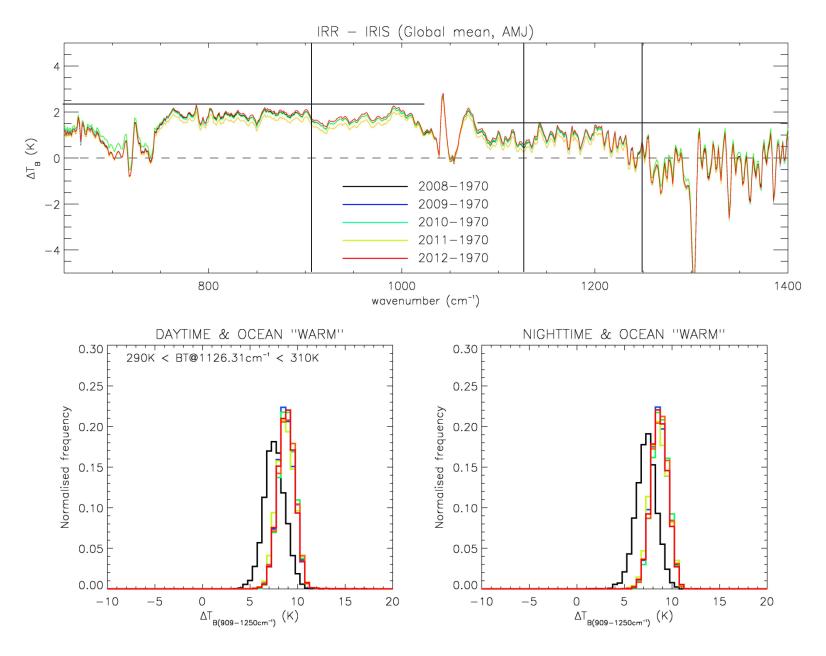


- Window inter-annual variability reduces most rapidly with increasing scale
   Simulations show the same behaviour but reduction in window is not as rapid. Non-window variability exceeds broadband at all scales and seems to show a faster rate of change with scale than observations
- Results in non-window variability becoming dominant at global scale
   Window variability still dominates at global scale
- Spectrally, global inter-annual variability < 0.17 K, < 0.05 K across window</li>
   Variability < 0.15 K but up to 0.08 K within window</li>

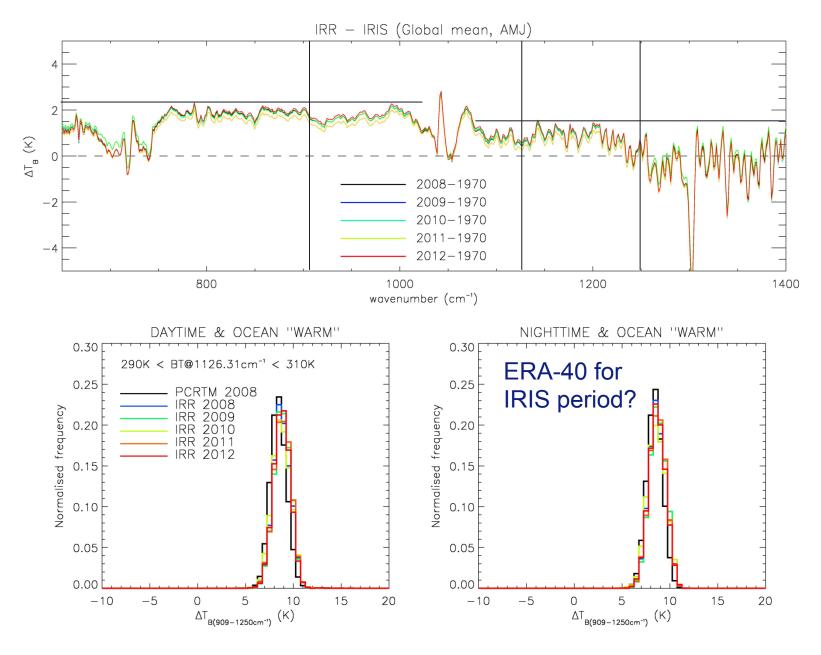
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Ocean only scenes, 220 K <  $T_{\rm B1126}$  < 250 K, IRR & IRIS

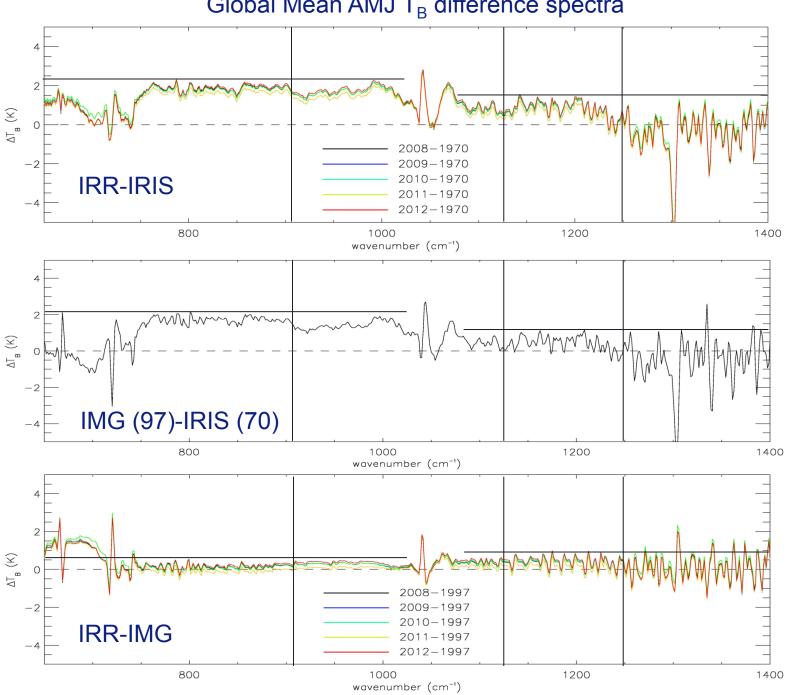


Ocean only scenes, 290 K < T<sub>B1126</sub> < 310 K, IRR & IRIS



Ocean only scenes, 290 K <  $T_{\rm B1126}$  < 310 K, IRR & PCRTM

#### Global Mean AMJ T<sub>B</sub> difference spectra



## Summary

- Used IASI data to probe how the emission to space varies spectrally on short timescales. While variability reduces with increasing spatial scale across the spectrum, the rate of change varies with wavenumber. Hence a more marked reduction is seen in window variability compared to that seen in regions sensitive to the upper troposphere.
- These findings are in agreement with observations from CERES over the same period and imply that at the largest spatial scales fluctuations in mid-upper tropospheric temperatures and water vapour, and not surface temperature or cloud, play the dominant role in determining the level of inter-annual all-sky OLR variability.
  - (Brindley et al., *J. Clim*, in review)
- Although simulations from reanalysis show an encouraging level of agreement in general, they do not replicate this scaling behaviour.
- To diagnose longer term spectral changes confidence in instrument calibration and stability is key.

#### TRUTHS/NCEO status

- New Minister for Universities and Science from July 2014 (Greg Clark): national/bilateral mission opportunities unlikely before UK general election.
- Presentations made to Mark Walport (Government Chief Scientific Advisor), lan Boyd (Chief Scientific Advisor at DEFRA), Outgoing DECC Chief Scientist. DECC Climate Minister visiting NPL in November.
- Joint bid with Airbus-DS for TRUTHS related funding in latest CEOI round (£1M budget). [Note that Imperial has also submitted a bid for FIR instrument development to same call].
- Formal letter of support for TRUTHS from NASA HQ?
- Ongoing CASE PhD studentship with Imperial to strengthen science case.
- Studies supporting CLARREO funded until at least April 2015 via NCEO (new director: John Remedios; HB now joint head of Earth Observation Data and Model evaluation division). Should be opportunities for continued funding for TRUTHS/CLARREO science support.